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## CaLCULATION OF THE FORCE AND MOMENT OF FORCES ACTING ON A DROP IN AN ARBITRARY NON-STEADY FLOW OF A VISCOUS FLUID*

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The Oseen point force method /1-3/ which differs from the methods used earlier in similar problems, is used to obtain formulas for the force and moment of forces acting on a spherical drop in an inhomogeneous nonsteady flow of viscous incompressible fluid. In special cases the results can be reduced to well-known results.

Earlier, the non-steady motion of a rigid particle in an inhomogeneous non-steady flow was considered in $/ 4,5 /$, its rotation in $/ 6,7 /$, the conditions of slippage at the surface in $/ 5,7 /$, and the effect of the specified external forces in $/ 8,9 /$. The corresponding stationary problem was studied in $/ 10,11 /$ and the non-steady motion of a drop in uniform non-steady flow in /12-14/.

1. Formulation of the problem. A liquid sphere of viscosity $\mu^{\prime}$, density $\rho^{\prime}$ and constant radius $a$, moves with velocity $u(t)$ through an incompressible medium of viscositv $\mu$ and density $\rho$. The problem is studied in the Stokes approximation, i.e. we consider the following linear, non-steady equations of motion of the fluid outside and inside the drop: đFrikl.Matcm.Mcklan., 50,5,772-779,1986

$$
\begin{gather*}
\rho \partial \mathbf{v} / \partial t=-\nabla p+\mu \Delta \mathbf{v}+f_{w}, \quad \operatorname{div} \mathbf{v}=0  \tag{1.1}\\
\rho^{\prime} \partial \mathbf{v}^{\prime} / \partial t=-\nabla p^{\prime}+\mu^{\prime} \Delta \mathbf{v}^{\prime}+\mathrm{f}_{w^{\prime}}, \quad \operatorname{div} \mathbf{v}^{\prime}=0 \tag{1.2}
\end{gather*}
$$

When there is no particles, the fluid has a velocity $\mathbf{v}_{\infty}(r, t)$ and a pressure $p_{\infty}(r, t)$, which also satisfy the system of Eqs. (1.1).

The conditions satisfied at the drop surface are the conditions that no fluid passes across the surface of contact and that the tangential components of the velocity are equal to the tangential components of the viscous stress tensor

$$
\begin{equation*}
v_{n}=v_{n}^{\prime}=u_{n}, \quad \mathbf{v}_{\tau}=\mathbf{v}_{\tau}^{\prime}, \quad \mathbf{p}_{n \tau}=\mathbf{p}_{n \tau}^{\prime} \tag{1.3}
\end{equation*}
$$

The following conditions hold away from the particle and at its centre:

$$
r \rightarrow \infty, \quad \mathbf{v}-\mathbf{v}_{\infty} \rightarrow 0, \quad p-p_{\infty} \rightarrow 0 ; \quad r=0, \quad\left|\mathbf{v}^{\prime}\right| \neq \infty, \quad p^{\prime} \neq \infty
$$

Let us introduce the dimensionless variables by dividing the radius vector, the time, velocity, pressure and the given external force by the quantities $a, a^{2} / v, v / a, \mu v / a a^{2}, \mu v / a^{3}(\mu=\rho \cdot v)$ respectively.

In the outer problem we can eliminate the given force $\mathrm{l}_{w}(\mathbf{r}, t)$ by subtracting the equations for $v_{\infty}, p_{\infty}$ from (1.1). We will assume that for the inner problem $f_{w}^{\prime}=\Gamma_{\varphi}$ (e.g. the force of gravity), and this can then be included in the pressure. Since the deviation of the drop shape from spherical (i.e. the conditions for $p_{n n}$ ) were not considered, it follows that the boundary conditions in this case remain unchanged.

As a result, we obtain the following equations and boundary conditions in terms of the dimensionless variables $\mathbf{w}=\mathbf{v}-\mathbf{v}_{\infty}, q=p-p_{\infty}, v^{\prime}, p^{\prime}$ :

$$
\begin{equation*}
\partial w / \partial t=-\nabla q+\Delta w, \operatorname{div} \mathbf{w}=0 \tag{1.4}
\end{equation*}
$$

$$
\begin{gather*}
\sigma B^{-2} \partial \mathbf{v}^{\prime} / \partial t=-\nabla p^{\prime}+\sigma \Delta \mathbf{v}^{\prime}, \operatorname{div} \mathbf{v}^{\prime}=0  \tag{1.5}\\
r=1, \quad w_{n}+v_{\propto n}=v_{n}^{\prime}=u_{n}, \quad \mathbf{w}_{\boldsymbol{\tau}}+\mathbf{v}_{\infty \tau}=\mathbf{v}_{\tau^{\prime}}^{\prime}  \tag{1.6}\\
\mathbf{q}_{n \boldsymbol{\tau}}+\mathbf{p}_{\propto n \tau}=\mathbf{p}_{n \tau}^{\prime} \\
r \rightarrow \infty, \mathbf{w} \rightarrow 0, q \rightarrow 0  \tag{1.7}\\
r=0,\left|\mathbf{v}^{\prime}\right| \neq \infty, p^{\prime} \neq \infty \tag{1.8}
\end{gather*}
$$

The system of Eqs. (1.4) and (1.5), which contains two dimensionless parameters $\sigma$ and $B^{2}$ equal to the ratios of the dynamic and kinematic coefficients inside and outside the drop respectively, together with conditions (1.6)-(1.8), enables us to determine the velocity and pressure field outside and inside the drop.
2. Constructing the solutions. we shall use the Laplace transform with parameter $s$, with respect to the argument of time $t$. Here the equations of motion take the form

$$
\begin{gather*}
s \mathbf{W}=-\nabla Q+\Delta \mathbf{W}, \operatorname{div} \mathbf{W}=0  \tag{2.1}\\
\sigma B^{-2} s \mathbf{V}^{\prime}=-\nabla P^{\prime}+\sigma \Delta \mathbf{V}^{\prime}, \quad \operatorname{div} \mathbf{V}^{\prime}=0  \tag{2.2}\\
r=1, \quad W_{n}+V_{\infty}=V_{n}^{\prime}=U_{n}, \quad \mathbf{W}_{\mathfrak{\imath}}+\mathbf{V}_{\infty \tau}=\mathbf{V}_{\tau}^{\prime},  \tag{2.3}\\
\mathbf{Q}_{n \mathfrak{r}}+\mathbf{P}_{\infty n \tau}=\mathbf{P}_{n \tau}^{\prime} \\
r \rightarrow \infty, \mathbf{W} \rightarrow 0, Q \rightarrow 0  \tag{2.4}\\
r=0, \quad\left|\mathbf{V}^{\prime}\right| \neq \infty, P^{\prime} \neq \infty \tag{2.5}
\end{gather*}
$$

Following the Oseen point force method /1, $2 /$, we shall construct the solutions of the above system of equations with help of the fundamental tensor $u_{k i}$ and vector $p_{k}$

$$
u_{k i}=\delta_{k i} \Delta \Phi-\frac{\partial^{2} \Phi}{\partial x_{k} \partial x_{i}}, \quad p_{k}=\frac{\partial}{\partial x_{k}}(s \mathrm{\Phi}-\Delta(\mathrm{D})
$$

The function $\Phi$ is a solution of the equation

$$
\Delta \Delta \Phi-s \Delta \Phi=\delta(\mathbf{r}), \quad \Phi=\left(1-e^{-\sqrt[V]{s} r}\right)(4 \jmath \stackrel{s}{ })^{-1}
$$

In this case, when $k$ is fixed the quantities $u_{k i}$ and $p_{k}$ satisfy the equations

$$
s \mathbf{u}_{k}-\nabla p_{k}+\Delta \mathbf{u}_{k}, \quad \operatorname{div} \mathbf{u}_{k}-0
$$

The solutions $F_{k}^{1} \mathbf{u}_{k}$ and $F_{k}^{1} p_{k}\left(F^{1}\right.$ is a constant vector) correspond to the flow generated by a point force situated at the origin of coordinates. Clearlv, any order derivatives with respect to the coordinations of $\mathbf{u}_{k}$ and $p_{k}$ will be solutions of system (2.1). In particular, we can say this of the quantities $v_{k i}, q_{k}$

$$
v_{k i}=\Delta u_{k i}=\delta_{k i} \Delta \Delta \Phi-\partial^{2} \Delta\left(\mathrm{D} / \partial x_{k} \partial x_{i}, \quad q_{k}=\Delta p_{k}=0 \quad(r \geqslant 1)\right.
$$

In order to satisfy the boundary conditions, we shall construct the outer solution of problem (2.1)-(2.5) with help of the derivatives with respect to the coordinates of the quantities $u_{k i}, v_{k i}$ and $p_{k}$. Collecting the terms accompanying $\Delta \Phi$ and separating in the corresponding
tensors the symmetric and antisymmetric parts, we obtain expressions for the velocity and pressure field, which in vector notation will be

$$
\begin{align*}
& \mathbf{W}=\boldsymbol{\Lambda} \Delta \Delta \Phi-\nabla(\boldsymbol{\Lambda} \cdot \nabla) \Delta \Phi+[\nabla \times \mathrm{L}] \Delta \Phi-\nabla(\mathbf{F} \cdot \nabla)\left(\frac{1}{4 . \pi r s}\right)  \tag{2.6}\\
& Q=(\mathbf{F} \cdot \nabla)\left(\frac{1}{4 \pi r}\right)
\end{align*}
$$

The components of the vectors $F$ and $L$ are linear differential operators of the form

$$
F^{i}=\sum_{n=1}^{\infty} F_{i j k \ldots l}^{(n)} \frac{\partial^{(n-1)}}{\partial x_{j} \partial x_{k} \ldots \partial x_{l}}, L^{i}=\sum_{n=1}^{\infty} L_{i j \ldots \ldots l}^{(n)} \frac{\partial^{(n-1)}}{\partial x_{j} \partial x_{k} \ldots \partial x_{l}}
$$

We can express the vector $\boldsymbol{\Lambda}$ in the same manner. We will show below that the boundary conditions at the sphere hold, if the tensors $F^{(n)}, L^{(n)}$ and $\Lambda^{(n)}$ are symmetrical over any pair of indices and the contractions over any two indices are equal to zero. The solution (2.6) represents another form of the general Lamb solution $/ 15 /$. This can be shown by assuming that quqntities of the form $f_{n}=\Lambda_{i j \ldots i}^{(n)} x_{i} x_{j} \ldots x_{k}$ are spherical volume harmonics of degree $n$, since they are homogeneous with respect to $r^{n}$ and $\Delta f_{n}=0$.

Indeed, we can write the solution (2.6) in the form

$$
\begin{align*}
\mathbf{W}= & \sum_{n=1}^{\infty}\left\{\left[s n^{-1} \Psi_{u-1}(r)-\psi_{n}(r)\right] \nabla\left(\Lambda_{i j \ldots h}^{(n)} x_{i} x_{i} x_{j} \ldots x_{k}\right)-\right.  \tag{2.7}\\
& r \Psi_{n+1}(r) \nabla r\left(\Lambda_{i j \ldots k}^{(n)} x_{i} x_{j} \ldots x_{i}\right)-n^{-1} \Psi_{n}(r) \operatorname{rot}\left[\mathbf{r}\left(L_{s q \ldots}^{(n)} l^{x} x_{s} x_{q} \ldots x_{l}\right)\right]- \\
& \left.\nabla\left[F_{l i j \ldots l}^{(n)} \frac{\partial^{(n)}}{\partial x_{k} \partial x_{j} \ldots \partial x_{l}}\left(\frac{1}{4 \pi s r}\right)\right]\right\}, Q=\sum_{n=1}^{\infty} F_{k j \ldots l}^{(n)} \frac{\partial^{(n)}}{\partial x_{k} \partial x_{j} \ldots \partial x_{l}}\left(\frac{1}{4 \pi r}\right)
\end{align*}
$$

The function $\psi_{n}(r)=\left(r^{-1} d / d r\right)^{n} \Delta \Phi$ satisfies the relations

$$
\frac{d^{2} \psi_{n}}{d r^{2}}+\frac{2(n+1)}{r} \frac{d \psi_{n}}{d r}-s \psi_{n}=0, \quad \frac{2 n+1}{r^{2}} \psi_{n}=\frac{s}{r^{2}} \psi_{n-1}-\psi_{n+1}
$$

For example, we have

$$
\begin{aligned}
& \psi_{0}(r)=\Delta \Phi=-\frac{1}{4 \pi r} e^{-\sqrt{s} r}, \psi_{1}(r)=\frac{1+\sqrt{s} r}{4 \pi r^{3}} e^{-\sqrt{s} r}, \\
& \psi_{1}(r)=-\frac{3+3 \sqrt{s r}+s r^{2}}{4 \pi r^{j}} e^{-\sqrt{s} r}
\end{aligned}
$$

The functions $\psi_{n}(r)$ are proportional to Bessel functions of fractional order / 16/. Remembering also that the tensors $F^{(n)}, L^{(n)}, A^{(n)}$ are symmetrical over any pair of indices and their contractions over two indices are equal to zero, i.e. that quantities of the form $f_{n}=$ $\Delta_{i, j}^{(n)} \ldots{ }_{k} x_{i} x_{j} \ldots x_{k}$ are spherical volume harmonics of degree $n$, we can establish a correspondence between the solution (2.6) and the general Lamb solution $/ 15 /$.

The solution of system (2.2) inside the drop, constructed in the same manner, has the form

$$
\begin{aligned}
& \mathbf{V}^{\prime}=-\nabla(S \cdot \nabla) \Delta()^{\prime}+S \Delta \Delta \Phi^{\prime}+[\nabla \times T] \Delta \Phi^{\prime}+ \\
& \quad \lambda^{-2} \nabla\left(\sum_{n=1}^{\infty} R_{i j \ldots l}^{(n)} x_{i} x_{j} \ldots x_{l}\right), \quad P^{\prime}=-\sigma \sum_{n=1}^{\infty} R_{i j \ldots l}^{(n)} x_{i} x_{j} \ldots x_{l} \\
& S^{i}=\sum_{n=1}^{\infty} S_{i j \ldots k}^{(n)} \frac{\partial^{(n-1)}}{\partial x_{j} \ldots \partial x_{l}}, \quad T^{i}=\sum_{n=1}^{\infty} T_{i j \ldots k}^{(n)} \frac{\partial^{(n-1)}}{\partial x_{j} \ldots \partial x_{k}}
\end{aligned}
$$

Here the tensors $S^{(n)}, T^{(n)}$ and $R^{(n)}$ are symmetrical over any pair of indices, and their contractions over two indices are equal to zero. The function $\Phi^{\prime}$ represents the solution of the equation

$$
\Delta \Delta \Phi^{\prime}-\lambda^{2} \Delta \Phi^{\prime}=0, \quad \Phi^{\prime}=\frac{\operatorname{sh} \lambda r}{r}, \quad \lambda^{2}=\frac{s}{B^{2}}, \quad \sigma=\frac{\mu^{\prime}}{\mu}, \quad B^{2}=\frac{v^{\prime}}{v}
$$

bounded at the point $r=0$.
We can establish, for the inner solution, a correspondence with the general Lamb solution, just as we did for the outer solution. Below, we shall make use of the stress vector, which can be found using the formula

$$
\begin{equation*}
Q_{n}=-Q_{\mathbf{n}}+\left(\frac{\partial}{\partial r}-\frac{1}{r}\right) \mathbf{W}+\frac{1}{r} \nabla(\mathbf{r} \cdot \mathbf{W})= \tag{2.8}
\end{equation*}
$$

$$
\begin{aligned}
& \sum_{k=1}^{\infty}\left\{2(k+2) r^{-1} \nabla\left(b_{k} r^{-k-1} F_{i j}^{(k)} \ldots l n_{i} n_{j} \ldots n_{l}\right)-\right. \\
& s\left(b_{k} r^{-k-1} F_{i j \ldots l}^{(k)}, n_{i} n_{j} \ldots n_{l}\right) \mathrm{n}-(k r)^{-1}\left[(k-1) \psi_{k}(r)+\right. \\
& \left.r^{2} \psi_{k+1}(r)\right] \operatorname{rot}\left[\mathrm{r}\left(L_{i}^{(i)} \ldots x_{i} x_{j} \ldots x_{l}\right)\right] \mid- \\
& (r k)^{-1}\left[\left(2 k^{2}-2+s r^{2}\right) \psi_{k}(r)-2 r^{2} \psi_{k_{i+1}}\right] \nabla\left(\Lambda_{i j \ldots i}^{(k)} x_{i} x_{j} \ldots x_{l}\right)+ \\
& \left.\left[2(k+2) \psi_{k+1}(r)-s \psi_{k}(r)\right]\left(\Lambda_{i j}^{(k)}, \ell x_{\imath} x_{j} \ldots x_{l}\right) \mathrm{n}\right\}, \\
& b_{k}=(-1)^{k}(2 k-1)!!/ 4 \pi s
\end{aligned}
$$

We can write the boundary conditions (2.3) in the form /10/

$$
\begin{gather*}
W_{n}+V_{\infty n}=V_{n}{ }^{\prime}=U_{n}, \quad \partial W_{n} / \partial r+\partial V_{\infty n} / \partial r=\partial V_{n}^{\prime} / \partial r  \tag{2.9}\\
(\mathbf{r} \cdot \operatorname{rot} \mathbf{W})+\left(\mathbf{r} \cdot \operatorname{rot} \mathbf{V}_{\alpha}\right)=\left(\mathbf{r} \cdot \operatorname{rot} \mathbf{V}^{\prime}\right),\left(\mathbf{r} \cdot \operatorname{rot} \mathrm{Q}_{n}\right)+\left(\mathbf{r} \cdot \operatorname{rot} \mathbf{P}_{\infty n n}\right)=\left(\mathbf{r} \cdot \operatorname{rot} \mathbf{P}_{n}{ }^{\prime}\right) \\
\left(\mathbf{r} \cdot \operatorname{rot}\left[\mathbf{r} \times \mathbf{Q}_{n}\right]\right)+\left(\mathbf{r} \cdot \operatorname{rot}\left[\mathbf{r} \times \mathbf{P}_{\infty n}\right]\right)=\left(\mathbf{r} \cdot \operatorname{rot}\left[\mathbf{r} \times \mathbf{P}_{n}^{\prime}\right]\right)
\end{gather*}
$$

We will represent the prescribed velocity field $\mathbf{V}_{\infty}(\mathbf{r}, s)$ in the form of a Taylor series

$$
V_{\infty}^{i}-U^{i}=\sum_{n=1}^{\infty} \frac{1}{(n-1)!}\left[\frac{\partial^{(n-1)}\left(V_{\infty}^{i}-U^{i}\right)}{\partial x_{j} \ldots \partial x_{p} \partial x_{q}}\right]_{0} x_{j} \ldots x_{p} x_{q}=\sum_{n=1}^{\infty} d_{i j \ldots p q}^{(i)} x_{j} \ldots x_{p} x_{q}
$$

The zero index means that the corresponding quantity is taken at the centre of the drop when the latter is not there.

We can expand the field $\mathbf{V}_{\infty}$ at the drop surface $r=1$ in terms of the independent spherical harmonics as follows:

$$
\begin{align*}
& V_{\infty n}-U_{n}=\sum_{k=1}^{\infty} d_{i j \ldots, q}^{(t)} n_{i} n_{j} \ldots n_{p} n_{q}=\sum_{k=1}^{\infty} A_{k}  \tag{2.10}\\
& \frac{\partial V_{\infty n}}{\partial r}=\sum_{k=1}^{\infty}(k-1) d_{i j \ldots p q}^{(k)} n_{i} n_{j} \ldots n_{p} n_{q}=\sum_{k=1}^{\infty} B_{k} \\
& -\left(r \cdot \operatorname{rot}\left[r \times \mathbf{P}_{\sim n}\right]\right)=\sum_{k=2}^{\infty}(k-1)\left[2(k+1) d_{i j \ldots}^{(k)}, n_{i} n_{j} \ldots n_{l}-\right. \\
& \left.(k-2) d_{i \eta}^{(i) \ldots} \ldots j n_{i} n_{j} \ldots n_{l}\right]=\sum_{k=1}^{\infty} E_{k} \\
& \left(\mathbf{r} \cdot \operatorname{rot} \mathbf{V}_{\alpha}\right)=\sum_{k=2}^{\infty}(k-2) \varepsilon_{i l j} d_{i j \neq \ldots q}^{(k)} n_{i} n_{p} \ldots n_{q}=\sum_{k=1}^{\infty} C_{k} \\
& \left(\mathbf{r} \cdot \operatorname{rot} \mathbf{P}_{\infty n}\right)=\sum_{k=2}^{\infty}(k-2)(k-1) \varepsilon_{i l j} d_{i j p \ldots q}^{(k)} n_{i} n_{p} \ldots n_{q}=\sum_{k=1}^{\infty} D_{k}
\end{align*}
$$

The quantities $A_{k}, B_{k}, E_{k}, C_{k}, D_{k}$ represent spherical surface harmonics of degree $k$. When $k=1$, the quantities can be easily calculated and used later in determining the force and the moment of forces acting on the drop. For example, we have for $A_{1}, B_{1}$

$$
\begin{aligned}
& A_{1}=\left(V_{\infty}^{i}-U^{i}\right) n_{i}+3 \sum_{k=1}^{\infty} \frac{\left[\Delta^{k} V_{\infty}{ }^{i}\right]_{0} n_{i}}{(2 k+3)(2 k+1)!}, \\
& B_{1}=3 \sum_{k=1}^{\infty} \frac{\left[\Delta^{k} V_{\infty}{ }^{i}\right]_{0} n_{i}}{(2 k+3)(2 k+1)(2 k-1)!}
\end{aligned}
$$

The boundary conditions (2.9) are reduced here to

$$
\begin{align*}
& (k+1) b_{k} F_{i j \ldots l}^{(k)} n_{i} n_{j} \ldots n_{l}+(k+1) \psi_{k} \Lambda_{i j}^{(k)} \ldots n_{i} n_{j} \ldots n_{l}+A_{k}=0  \tag{2.11}\\
& k R_{i j \ldots l}^{(k)} n_{i} n_{j} \ldots n_{l}+\lambda^{2}\left(k+1 ; p_{k}{ }^{\prime} S_{i j \ldots l}^{(i)} n_{i} n_{j} \ldots n_{l}=0\right. \\
& (k+1)\left\{-(k+2) b_{k} F_{j}^{(k)} \ldots l n_{i} n_{j} \ldots n_{l}+\right. \\
& \left.\left[(k-1) \psi_{k}+\psi_{k+1}\right] \Lambda_{i j}^{(!)} \ldots n_{1} n_{j} \ldots n_{l}\right]+B_{k}=k(k-1)^{\lambda^{-2}} R_{i j}^{\prime(k)} \ldots n_{i} n_{j} \ldots n_{l}+ \\
& (k+1)\left[\psi_{k+1}^{\prime}+(k-1) \psi_{k}{ }^{\prime}\right] S_{i j \ldots i}^{(k)} n_{i} n_{j} \ldots n_{l} \\
& 2 b_{k} k(k+2) F_{i j}^{(h)} . . n_{i} n_{j} \ldots n_{l}+(k+1)^{-1} E_{k}+-
\end{align*}
$$

$$
\begin{aligned}
& \left.\left[\left(2 k^{2}-2+\lambda^{2}\right) \psi_{k}^{\prime}-2 \psi_{i+1}^{\prime}\right] S_{i j \ldots i}^{(t)} n_{i} n_{j} \ldots n_{l}\right\}
\end{aligned}
$$

$$
\begin{aligned}
& (k+1) \psi_{k} L_{i j \ldots}^{(\cdot)} n_{i} n_{j} \ldots n_{l}+C_{k}=(k+1) \psi_{k}^{\prime} T_{i j}^{(l)} . . n_{i} n_{j} \ldots n_{l} \\
& (k+1)\left[(k-1) \psi_{k}+\psi_{k-1}\right] L_{i j \ldots}^{(j)} n_{i} n_{j} \ldots n_{l}+D_{k}= \\
& \sigma(k+1)\left[(k-1) \psi_{i i}^{\prime}+\psi_{k-1}^{\prime}\right] T_{i j \ldots l}^{\left(b_{j}\right)} n_{i} n_{j} \ldots n_{l}
\end{aligned}
$$

where

$$
\begin{aligned}
& \psi_{k}=\left\{\psi_{k}(r)\right\}_{r=1}, \quad \psi_{k}{ }^{\prime}=\left\{\left(r^{-1} d / d r\right)^{k} \Delta \Phi^{\prime}\right\}_{r=1}, \\
& R_{i j \ldots l}^{(k)}=R_{i j \ldots l}^{(k)}(k \neq 1), \quad R_{i}^{(1)}=R_{i}^{(\mathbf{1})}-U^{i}
\end{aligned}
$$

The functions $\psi_{k}^{\prime}(r)$ satisfy relations analogous to the equations for $\psi_{k}(r)$.
The solutions of the system of Eqs.(2.11) are as follows:

$$
\begin{align*}
& \Lambda_{i j \ldots, \ldots}^{(1)} n_{i} n_{j} \ldots n_{l}=\left\{(k+2)\left[2 k \psi_{k+1}^{\prime}-\sigma\left(2 \psi_{k+1}^{\prime}-\lambda^{2} \psi_{k}{ }^{\prime}\right)\right] A_{k}-\right.  \tag{2.12}\\
& \left.\psi_{i+1}^{\prime} E_{k}-\sigma\left(2 \psi_{k+1}^{\prime}-\lambda^{2} \psi_{k}^{\prime}\right) B_{k}\right\}\left(s \chi_{k}\right)^{-1} \\
& S_{i j \ldots l}^{(k)} n_{i} n_{j} \ldots n_{l}=\left\{(k \mid 2)\left[2(k-1) \psi_{k-1} \mid \psi_{k}\right] A_{k}\right. \\
& \left.\psi_{i-1} E_{k}+\left(\psi_{k}-2 \psi_{k-1}\right) B_{k}\right\} \chi_{k}^{-1} \\
& F_{i j \ldots l}^{(k)} n_{i} n_{j} \ldots n_{l}=-\left[\psi_{k} \Lambda_{i j \ldots l}^{(k)}, n_{i} n_{j} \ldots n_{l}+(k+1)^{-1} A_{k}\right] b_{k}^{-1} \\
& R_{i j \ldots}^{\prime(k)} n_{i} n_{j} \ldots n_{l}=-(k+1) k^{-1} \lambda^{2} \psi_{k}^{\prime} S_{i j \ldots l}^{(i)} n_{i} n_{j} \ldots n_{l} \\
& L_{i j}^{( }{ }^{\prime}, n_{i} n_{i} n_{j} \ldots n_{l}=\left\{\psi_{k}^{\prime} D_{k}-\sigma\left[(k-1) \psi_{k}^{\prime}+\psi_{i+1}^{\prime}\right] C_{k}\right\} \varphi_{i}^{-1} \\
& T_{i j \ldots}^{(i)}, n_{i} n_{j} \ldots n_{l}=\left\{\psi_{k} D_{k}-\sigma\left[(k-1) \psi_{k}+\psi_{k+1}\right] C_{k}\right\} \varphi_{k}^{-1} \\
& \chi_{k}(s)=(k+1)\left[\sigma\left(2 \psi_{k+1}^{\prime}-\lambda^{2} \psi_{k}{ }^{\prime}\right) \psi_{k-1}+\psi_{k+1}^{\prime}\left(\psi_{k}-2 \psi_{i-1}\right)\right] \\
& \varphi_{k}(s)=(k+1)\left\{\sigma\left[(k-1) \psi_{k}^{\prime}+\psi_{k+1}^{\prime}\right] \psi_{k}-\left[(k-1) \psi_{k}+\psi_{k+1}\right] \psi_{k}{ }^{\prime}\right\}
\end{align*}
$$

As we have already said, the quantities $A_{k}, B_{k}, C_{k}, D_{k}, E_{k}$ are known, independent spherical surface harmonics of degree $k$, expressed in terms of the given field $\mathbf{V}_{\infty}$ by the formulas (2.10). They are easily computed for $k=1$, and appear in the expressions for the force and moment of forces which are found from the surface integrals containing the stress tensor (2.8) and depending, by virtue of the orthogonal nature of the harmonics, only on the vectors $\mathbf{F}^{1}$ and $\mathbf{L}^{1}$ shown in the solutions (2.12).
3. Formulas for the force and moment of forces. The quantities shown are related to the vectors $\mathbf{F}^{1}$ and $\mathbf{L}^{1}$ which correspond to a point force and point moment and are determined from Eqs. (2.11) by the relations

$$
\begin{aligned}
& \mathbf{G}(s)=\mathbf{F}^{\mathbf{1}}(s)+4 / 3 \pi s \mathbf{U}(s)-\int \mathbf{F}_{w}(\mathbf{r}, s) d \mathbf{r} \\
& \mathbf{M}(s)={ }^{2} / 3(3+3 \sqrt{s}+s) e^{-1 / 5} \mathbf{L}^{\mathbf{1}}+s \int\left[\mathbf{r} \times \mathbf{V}_{\infty}\right] d \mathbf{r}-\int\left[\mathbf{r} \times \mathbf{F}_{w}\right] d \mathbf{r}
\end{aligned}
$$

Here the integration is carried out over the volume of the unit sphere. Using the inverse Laplace transform, we finally obtain the following expressions for the force $g(t)$ and moment of forces $\mathbf{m}(t)$ acting on the drop in non-uniform non-steady flow:

$$
\begin{aligned}
& \mathrm{g}(t)=2 \pi \frac{33+9}{s+1}\left(\left[\mathbf{v}_{\alpha}\right]_{0}-\mathbf{u}\right)+6 \int_{0}^{t} d_{\mathrm{\tau}} I_{1}(t-\mathrm{\tau}) \frac{d}{d \tau}\left(\left[\mathbf{v}_{\star}\right]_{0}-\mathbf{u}\right) \div \\
& \left.\left.\frac{2}{3} \pi \frac{d}{d t}\left(\mid v_{x}\right]_{u}-u\right): \left.\frac{4}{3} \pi \frac{d}{d t} \right\rvert\, \mathbf{v}_{\mathbf{x}}\right]_{0}+ \\
& \sum_{k=1}^{\infty} \frac{1}{(2 k+1)!}\left\{2 \pi \frac{3 J+2(1-k)}{5+1}\left[\Delta^{k} \mathbf{v}_{\infty}\right]_{0}+\frac{6 \pi}{2 k+3} \frac{d}{d t}\left[\Delta^{k} \mathbf{v}_{\infty}\right]_{0}+\right. \\
& \left.6 \int_{0}^{t} d \tau\left[(1-k) I_{1}(t-\tau)+k I_{2}(t-\tau)\right] \frac{d}{d \tau}\left[\Delta^{k} \mathbf{v}_{\infty}\right]_{0}\right\}-\int \mathbf{f}_{w}(\mathbf{r}, t) d \mathbf{r} \\
& \mathbf{m}(t)=\frac{4}{3} \int_{0}^{t} d \tau I_{3}(t-\tau) \frac{d}{d \tau}\left[\operatorname{rot} \mathbf{v}_{\propto}\right]_{0}+\frac{4}{15} \pi \frac{d}{d t}\left[\operatorname{rot} \mathbf{v}_{\infty}\right]_{0}+ \\
& \sum_{k=1}^{\infty} \frac{4}{(\overline{2 k+3)}(2 k+1)!}\left\{-2 \pi k\left[\mathbf{r o t} \Delta^{k} \mathbf{v}_{\infty}\right]_{0}+\frac{\pi}{2 k+5} \frac{d}{d t}\left[\mathbf{r o t} \Delta^{k} \mathbf{v}_{\infty}\right]_{0}+\right. \\
& \left.\int_{0_{i}}^{t} d \tau\left[I_{3}(t-\tau) \div 2 k I_{4}(t-\tau)\right] \frac{d}{d \tau}\left[\boldsymbol{r o t} \Delta^{k} \mathbf{v}_{\alpha}\right]_{0}\right\}-\int\left[\mathbf{r} \times \mathbf{f}_{w}\right] d \mathbf{r}
\end{aligned}
$$

The Laplace transforms of the functions $I_{i}(t)$ are, respectively, $\pi s^{-1} K_{i}(s)(i=1,2,3,4)$

$$
\begin{aligned}
& K_{1}(s)=\xi(s)[\sigma \alpha(\lambda)+\beta(\lambda)]-(3 \sigma+2)(3 \sigma+3)^{-1}, \\
& K_{2}(s)=\sigma \alpha(\lambda) \xi(s)-\sigma(\sigma+1)^{-1} \\
& K_{3}(s)=(3+3 \sqrt{s}+s) \theta(s), \quad K_{4}(s)=(1+\sqrt{s}) \theta(s) \\
& \xi(s)=(1+\sqrt{s})[\sigma \alpha(\lambda)+(3+\sqrt{s}) \beta(\lambda)]^{-1}, \\
& \theta(s)=\sigma \lambda^{2} \beta(\lambda)\left[\sigma \lambda^{2}(1+\sqrt{s}) \beta(\lambda)+(3+3 \sqrt{s}+s) \gamma(\lambda)\right]^{-1} \\
& \gamma(\lambda)=\lambda \operatorname{ch} \lambda-\operatorname{sh} \lambda, \beta(\lambda)=\operatorname{sh} \lambda-3 \gamma(\lambda) / \lambda^{2}, \\
& \alpha(\lambda)=\gamma(\lambda)-2 \beta(\lambda)
\end{aligned}
$$

When the particle is rigid, we obtain

$$
\begin{aligned}
& I_{1}(t)=I_{2}(t)=\frac{\sqrt{\pi}}{\sqrt{t}}, \quad I_{3}(t)=\pi\left(3-e^{i} \operatorname{erfc} \sqrt{\bar{t}}+\frac{1}{\sqrt{\pi t}}\right), \\
& I_{4}(t)=\pi
\end{aligned}
$$

and in the case of a bubble we have, as $\sigma \rightarrow 0$,

$$
I_{1}(t)=4 / 3 \pi e^{8 t} \operatorname{erfc}(3 \sqrt{t}), I_{2}(t)=I_{3}(t)=I_{4}(t)=0
$$

These two special cases agree with the results obtained in $/ 4-9 /$. When $\sigma=B=1$, we can write the following analytic expressions for $I_{3}(t), I_{4}(t)$ :

$$
\begin{aligned}
& I_{3}(t)=\frac{\sqrt{\pi}}{2 \sqrt{t}}\left[12 t^{2}-6 t+1-\left(12 t^{2}+6 t+1\right) \exp \left(-\frac{1}{t}\right)\right] \\
& I_{4}(t)=\pi\left[\frac{2 t^{2 / 2}}{\sqrt{\pi}}-\frac{2 \sqrt{t}}{\sqrt{\pi}}+\frac{1}{2}-\frac{2 t^{2 / 2}}{\sqrt{\pi}} \exp \left(-\frac{1}{t}\right)-\frac{1}{2} \operatorname{erfc}\left(\frac{1}{\sqrt{t}}\right)\right]
\end{aligned}
$$

For other values of the parameters $\sigma$ and $B$ the quantities $I_{i}(t)$ are determined numerially. The results of the computations are shown in Figs.1-6. Curves $1-6$ correspond to values of the parameters $B^{2}$ equal to $0.001 ; 0.01 ; 0.1 ; 1 ; 10 ; 100$.

For short times, the following asymptotic formulas hold:

$$
\begin{aligned}
& I_{1}(t)=\frac{\sigma \sqrt{\pi}}{(\sigma+B) \sqrt{\bar{t}}}+\frac{\sigma^{2}+2 b^{2}}{(\sigma+B)^{2}} \pi-\frac{3 J+2}{3(5+1)} \pi+\frac{2 B\left[3 \sigma B^{2}(\sigma+B)-(\sigma-2 B)^{2}\right] \sqrt{\pi t}}{(\sigma+B)^{3}}+\ldots \\
& I_{2}(t)=\frac{\sigma \sqrt{\pi}}{(\sigma+B) \sqrt{l}}+\frac{\sigma(\sigma-2 B)}{(\sigma+B)^{2}} \pi+\frac{\sigma}{\sigma+1} \pi+\frac{\hat{\beta}^{2} \mathrm{~B}\left(B^{2}+B J+2-\sigma\right) \sqrt{\pi} \bar{t}}{(\sigma+B)^{3}}+\ldots \\
& I_{3}(t)=\frac{\sigma \sqrt{\pi}}{(5+B) \sqrt{t}}+\frac{2 \pi J}{(5+B)^{2}}\left(\sigma-B^{2}\right)+\ldots, \quad I_{4}(t)=\frac{5}{5-B} \pi-\frac{\left.45 B_{1} B+1\right)}{\left(5+B,^{2}\right.} \sqrt{\pi t}+\ldots
\end{aligned}
$$



Fig. 1


Fig. 3


Fig. 2


Fig. 4


As $t \rightarrow \infty$, the functions $I_{i}(t)$ tend to

$$
\begin{aligned}
& I_{1}(t)=\frac{(2+3 s)^{2}}{9(1+\sigma)^{2}} \sqrt{\pi}-\frac{2+3 s}{\sqrt{t}}\left[\frac{2 \sigma}{54(1+\sigma)^{3}}+\frac{2+3 s}{3(1+\sigma)}\right] \frac{\sqrt{\pi}}{t^{3 / 2}}+\ldots, \\
& I_{2}(t)=\frac{\sigma(2+35)}{3(1+\sigma)^{2}} \frac{\sqrt{\pi}}{\sqrt{t}}-\frac{\sigma}{18(1+\sigma)^{3}}\left[\frac{4 s+2}{7 B^{2}}+\frac{2+3 s}{3(1+\sigma)}\right] \frac{\sqrt{\pi}}{t^{3 / 2}}+\ldots, \\
& I_{3}(t)=\frac{\sigma^{2}}{120 B^{3}} \frac{\sqrt{\pi}}{t^{1 / 2}}+\ldots \\
& I_{4}(t)=\frac{\sigma}{60 B^{2}} \frac{\sqrt{\pi}}{t^{3 / 2}}+\ldots
\end{aligned}
$$

The ends of the curves infigs.1-6 are described by the asymptotic formulas with an error not exceeding $10-15 \%$.

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